

# A Plumblines Constraint for the Rational Function Lens Distortion Model

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Calibration of lens distortion has traditionally proceeded in one of three ways. The first is to use known correspondences between feature points in one or more images, and the world 3D points. This is typically done using a checkerboard or a calibration grid of dots where corners or dot centres can be reliably located [7]. A second class of calibration techniques is termed auto-calibration as it relies solely on detecting static points within a scene [4, 8]. This paper fits into the third category, where straight lines in the world are used to determine the distortion parameters. This *plumblines* technique was first mooted by [1], and has been applied to various distortion models [3, 6].

We show in this paper how a plumblines constraint can be implemented using the rational function model for lens distortion [2], under which straight lines are imaged as conics, and this permits an elegant factorization of the conics into the camera calibration and the equations of the straight lines. This differs from previous plumblines work in two ways: first, a factorization-based algorithm can be formulated to estimate the distortion; second, nonlinear refinement of the distortion can be easily implemented to minimize a good approximation of geometric distance in the image plane. While this was possible with previous models, the simplicity of the mapping in this case appears to lead to fast and efficient convergence of the nonlinear algorithm over a range of starting positions.

For a perspective camera, the mapping from image pixels  $(i, j)$  to 3D rays  $\mathbf{d}(i, j)$  can be expressed as:

$$\mathbf{d}(i, j) = \begin{pmatrix} B_{11}i + B_{12}j + B_{13} \\ B_{21}i + B_{22}j + B_{23} \\ B_{31}i + B_{32}j + B_{33} \end{pmatrix} = \mathbf{B} \begin{pmatrix} i \\ j \\ 1 \end{pmatrix}, \quad (1)$$

where the  $3 \times 3$  matrix  $\mathbf{B} = \mathbf{R}^\top \mathbf{K}^{-1}$ , and  $\mathbf{R}$  is often chosen to be the identity [5]. The rational function model handles lens distortion by permitting  $i$  and  $j$  to appear in higher order polynomials, in particular quadratic:

$$\mathbf{d}(i, j) = \begin{pmatrix} A_{11}i^2 + A_{12}ij + A_{13}j^2 + A_{14}i + A_{15}j + A_{16} \\ A_{21}i^2 + A_{22}ij + A_{23}j^2 + A_{24}i + A_{25}j + A_{26} \\ A_{31}i^2 + A_{32}ij + A_{33}j^2 + A_{34}i + A_{35}j + A_{36} \end{pmatrix}. \quad (2)$$

This model may be written as a linear combination of the distortion parameters, in a  $3 \times 6$  matrix  $\mathbf{A}$  (analogous to  $\mathbf{B}$  above), and a 6-vector  $\mathcal{X}$  of monomials in  $i$  and  $j$ . Define  $\mathcal{X}$  as the “lifting” of image point  $(i, j)$  to a six dimensional space

$$\mathcal{X}(i, j) = [i^2, ij, j^2, i, j, 1]^\top \quad (3)$$

The imaging model (2) may then be written

$$\mathbf{d}(i, j) = \mathbf{A}\mathcal{X}(i, j) \quad (4)$$

A line in the scene forms a plane with the origin of camera coordinates, and is imaged to the set of  $\mathbf{d}$  in that plane. This yields the line equation  $\mathbf{1}^\top \mathbf{d} = 0$  which, in terms of image points  $(i, j)$  becomes

$$\mathbf{1}^\top \mathbf{A}\mathcal{X} = 0 \iff \boldsymbol{\theta}^\top \mathcal{X} = 0, \quad (5)$$

where  $\boldsymbol{\theta} = (A_{11}, A_{12}, A_{13}, A_{14}, A_{15}, A_{16})^\top$  are the parameters of a conic in image coordinates  $(i, j)$  and  $\mathcal{X}$  is given by (3). Here we observe the important property that lines in the world go to conics under the rational function model. The task of calibration is then to find an  $\mathbf{A}$  which will map these conics in the distorted image back to straight lines.

By fitting a conic to the image of the line, we obtain parameters  $\boldsymbol{\theta}$ , and thus the constraint

$$\boldsymbol{\theta} = \mathbf{A}^\top \mathbf{1}$$

for unknown  $\mathbf{A}$  and  $\mathbf{1}$ . The equality is exact as any scale factor is included in  $\mathbf{1}$ . Collecting  $L$  such constraints, we obtain

$$\underbrace{[\boldsymbol{\theta}_1 \mid \dots \mid \boldsymbol{\theta}_L]}_{6 \times L} = \underbrace{\mathbf{A}^\top}_{6 \times 3} \underbrace{[\mathbf{1}_1 \mid \dots \mid \mathbf{1}_L]}_{3 \times L}$$

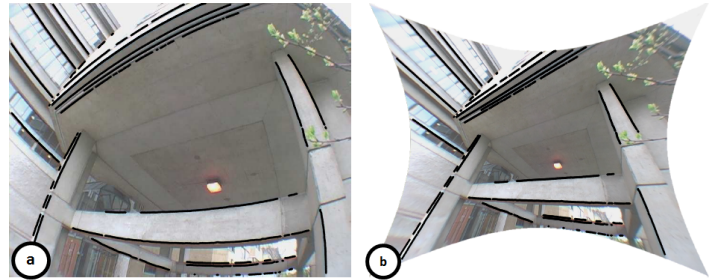


Figure 1: Edges corresponding to straight lines in the real world are detected (a) and the plumblines constraint is used to compute the distortion parameters, giving the rectified image (b)

which we write as

$$\mathbf{C} = \mathbf{A}^\top \mathbf{L}$$

so the matrix of conic parameters  $\mathbf{C}$  is of rank no greater than 3. Therefore  $\mathbf{A}$  can be computed up to a homography by factorization: if  $\mathbf{U}\mathbf{S}\mathbf{V}^\top = \mathbf{C}$  is the SVD of  $\mathbf{C}$ , then  $\mathbf{A} = \mathbf{S}_{(1:3,1:3)}\mathbf{U}_{(:,1:3)}^\top$  is one member of the equivalence class of solutions.

The matrix  $\mathbf{C}$  will not be rank 3 if the conics were obtained by fitting to noisy image data. The above factorization truncates  $\mathbf{C}$  to rank 3 by minimizing the error in the conic parameter space, and as discussed in the paper, this is sensitive to noise. The strategy for fitting  $\mathbf{A}$  from noisy image data is to run a non-linear optimization that finds the  $\mathbf{A}$  which minimizes the error between the image data and straight lines projected (as conics) into the distorted image. The nonlinear objective measures the Sampson distance [5] from the conics to the detected edges. Let  $\mathbf{e}_{\ell k} = (i_{\ell k}, j_{\ell k})$  denote the  $k^{\text{th}}$  edgel in linked segment  $\ell$ . The Sampson distance is a first order approximation to the distance from a point to a conic. The error function  $\mathcal{E}(\mathbf{A}, \mathbf{l}_1, \dots, \mathbf{l}_L)$  we minimize is then given by (with  $\boldsymbol{\theta}^\ell := \mathbf{A}\mathbf{l}_\ell$ )

$$\mathcal{E} = \sum_{\ell=1}^L \sum_{k=1}^{n_\ell} \frac{[(\boldsymbol{\theta}^\ell)^\top \mathcal{X}(i_{\ell k}, j_{\ell k})]^2}{(2\theta_1^\ell i_{\ell k} + \theta_2^\ell j_{\ell k} + \theta_4^\ell)^2 + (2\theta_3^\ell j_{\ell k} + \theta_2^\ell i_{\ell k} + \theta_5^\ell)^2}$$

Implementation of this method by edge detection and edge linking is described in the paper, as are the details of the nonlinear optimization. Our conclusion is that the simplicity of the rational function model, coupled with its ability to model a variety of lenses, makes it a useful model to consider when dealing with lenses exhibiting moderate to severe distortion.

- [1] D. C. Brown. Close-range camera calibration. *Photogrammetric Eng.*, 37(8):855–866, 1971.
- [2] D. Claus and A. W. Fitzgibbon. A rational function lens distortion model for general cameras. In *Proc. CVPR*, pages 213–219, 2005.
- [3] F. Devernay and O. Faugeras. Straight lines have to be straight. *MVA*, 13:14–24, 2001.
- [4] A. W. Fitzgibbon. Simultaneous linear estimation of multiple view geometry and lens distortion. In *Proc. CVPR*, 2001.
- [5] R. I. Hartley and A. Zisserman. *Multiple View Geometry in Computer Vision*. Cambridge University Press, ISBN: 0521623049, 2000.
- [6] R. Swaminathan and S. Nayar. Nonmetric calibration of wide-angle lenses and polycameras. *IEEE T-PAMI*, 22(10):1172–1178, 2000.
- [7] Y. R. Tsai. An efficient and accurate camera calibration technique for 3D machine vision. In *Proc. CVPR*, 1986.
- [8] Z. Zhang. On the epipolar geometry between two images with lens distortion. In *Proc. ICPR*, pages 407–411, 1996.